LECTURE 29: SPHERICAL (II) + THE JACOBIAN (I)

1. More Spherical Practice

Example 1:

$$\int \int \int_E x \, dx dy dz$$

E: solid under the cone $z = \sqrt{x^2 + y^2}$ and inside the sphere $x^2 + y^2 + z^2 = 1$, in the first octant.

STEP 1: Picture:



Date: Wednesday, November 3, 2021.

Note:
$$z = \sqrt{x^2 + y^2} \Rightarrow \phi = \frac{\pi}{4}$$
 and $x^2 + y^2 + z^2 = 1 \Rightarrow \rho = 1$

STEP 2: Inequalities: (here we have $\frac{\pi}{2}$ because it's the first octant)

$$\begin{cases} 0 \le \rho \le 1\\ 0 \le \theta \le \frac{\pi}{2}\\ \frac{\pi}{4} \le \phi \le \frac{\pi}{2} \end{cases}$$

STEP 3: Integrate:

$$\begin{split} &\int \int \int_{E} x \, dx \, dy \, dz \\ &= \int_{\frac{\pi}{4}}^{\frac{\pi}{2}} \int_{0}^{\frac{\pi}{2}} \int_{0}^{1} \rho \sin(\phi) \cos(\theta) \, \rho^{2} \sin(\phi) \, d\rho \, d\theta \, d\phi \\ &= \left(\int_{0}^{1} \rho^{3} d\rho \right) \left(\int_{0}^{\frac{\pi}{2}} \cos(\theta) \, d\theta \right) \left(\int_{\frac{\pi}{4}}^{\frac{\pi}{2}} \sin^{2}(\phi) \, d\phi \right) \\ &= \left[\frac{\rho^{4}}{4} \right]_{0}^{1} \left[\sin(\theta) \right]_{0}^{\frac{\pi}{2}} \int_{\frac{\pi}{4}}^{\frac{\pi}{2}} \frac{1}{2} - \frac{1}{2} \cos(2\phi) \, d\phi \\ &= \left(\frac{1}{4} \right) \left(1 \right) \left[\frac{\phi}{2} - \frac{1}{4} \sin(2\phi) \right]_{\frac{\pi}{4}}^{\frac{\pi}{2}} \left(u = 2\phi \right) \\ &= \frac{1}{4} \left(\frac{\pi}{4} - \frac{\pi}{8} - \frac{1}{4} \sin(\pi) + \frac{1}{4} \sin\left(\frac{\pi}{2}\right) \right) \\ &= \frac{1}{4} \left(\frac{\pi}{8} + \frac{1}{4} \right) \\ &= \frac{\pi}{32} + \frac{1}{16} \end{split}$$

2. Mass of the sun

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Video: Mass of the Sun
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And of course I saved the best for last! Because there is this meme that was popular a couple of years ago:



Without further ado ... let's calculate the mass of the sun!

Example 2:

Suppose the sun E is a ball of radius $R = 6.9 \times 10^{10}$ cm and density $\frac{1}{\sqrt{x^2+y^2+z^2}}$ g/cm^3 . What is the mass of the sun?

Note: $\frac{1}{\sqrt{x^2+y^2+z^2}}$ blows up near (0,0,0), so this is saying that the sun is heavier at its core than on the surface, which makes sense physically.

STEP 1: Picture:



STEP 2: Inequalities:

$$\begin{cases} 0 \le \rho \le R\\ 0 \le \theta \le 2\pi\\ 0 \le \phi \le \pi \end{cases}$$

Mass
$$= \int \int \int_{E} \frac{1}{\sqrt{x^2 + y^2 + z^2}} dx dy dz$$
$$= \int_{0}^{\pi} \int_{0}^{2\pi} \int_{0}^{R} \left(\frac{1}{\rho}\right) \rho^2 \sin(\phi) d\rho d\theta d\phi$$
$$= 2\pi \left(\int_{0}^{R} \rho d\rho\right) \left(\int_{0}^{\pi} \sin(\phi) d\phi\right)$$
$$= (2\pi) \frac{R^2}{2} (2)$$
$$= 2\pi R^2$$
$$\approx 2.99 \times 10^{22} g$$

Remark: NASA uses the following density:

$$519\left(\frac{\rho}{R}\right)^4 - 1630\left(\frac{\rho}{R}\right)^3 + 1844\left(\frac{\rho}{R}\right)^2 - 889\left(\frac{\rho}{R}\right) + 155$$

Which would give you 2.7×10^{33} grams. The actual mass is 1.98×10^{33} grams, so this is not bad at all!

3. u-sub the COOL way

Welcome to the one and only integration technique in this course: u-sub! For this, let me "remind" you how to do single-variable u-sub, but I'll present it in a way that will be useful in this course

Example 3:

Calculate $\int_1^2 e^{-x^2}(-2x)dx$

STEP 1: Let $u = -x^2$

STEP 2: Endpoints: u(1) = -1, u(2) = -4.

So *u* turns D = [1, 2] into D' = [-1, -4] = [-4, -1].



STEP 3: du: Beware of the absolute value! (makes sense, du should be positive)

$$du = \left| \frac{du}{dx} \right| dx = \left| -2x \right| dx = 2x dx \Rightarrow -2x dx = -du$$

STEP 4: Integrate

$$\int_{1}^{2} e^{-x^{2}}(-2x)dx = \int_{[1,2]} e^{-x^{2}}(-2x)dx = \int_{D} e^{-x^{2}}(-2x)dx = \int_{D'} e^{u}(-du)$$
$$= -\int_{[-4,-1]} e^{u}du = -\int_{-4}^{-1} e^{u}du = -(e^{-1} - e^{-4}) = e^{-4} - e^{-1}$$

4. Multivariable Example

Video: The Jacobian

The good news is that for double and triple integrals, the process is similar to the above!

Example 4:

$$\int \int_D \sin\left(\frac{y-x}{y+x}\right) \, dx \, dy$$

Where D is the square with vertices (-1, 0), (0, -1), (1, 0), (0, 1).

STEP 1:

$$\begin{cases} u = y - x \\ v = y + x \end{cases}$$

STEP 2: "Endpoints"

Trick: Look at the values of u and v at the vertices:



$$(-1,0) \Rightarrow \begin{cases} u = y - x = 0 - (-1) = 1\\ v = y + x = 0 + (-1) = -1 \end{cases} \Rightarrow (1,-1)$$
$$(0,-1) \Rightarrow \begin{cases} u = -1 - 0 = -1\\ v = -1 + 0 = -1 \end{cases} \Rightarrow (-1,-1)$$

Similarly (1,0) becomes (-1,1) and (0,1) becomes (1,1).

So D' is a square with vertices (1, -1), (-1, -1), (-1, 1), (1, 1)STEP 3: " $du = \left|\frac{du}{dx}\right| dx''$

Here we get:
$$dudv = \left| \frac{dudv}{dxdy} \right| dxdy$$

Let's put all the possible partial derivatives together in a determinant:

$$\frac{dudv}{dxdy} = \begin{vmatrix} \frac{\partial u}{\partial x} & \frac{\partial u}{\partial y} \\ \frac{\partial v}{\partial x} & \frac{\partial v}{\partial y} \end{vmatrix} = \begin{vmatrix} -1 & 1 \\ 1 & 1 \end{vmatrix} = (-1)(1) - (1)(1) = -2$$

Therefore:
$$dudv = |-2| dxdy = 2dxdy \Rightarrow dxdy = \frac{1}{2} dudv$$

Note: This number (or its absolute value) is called the Jacobian, a tribute to Taylor Lautner in Twilight. It's sometimes written as $\frac{\partial(u,v)}{\partial(x,y)}$ instead of $\frac{dudv}{dxdy}$



STEP 4: Integrate:

$$\int \int_{D} \sin\left(\frac{y-x}{y+x}\right) dx dy = \int \int_{D'} \sin\left(\frac{u}{v}\right) \frac{1}{2} du dv$$
$$= \frac{1}{2} \int_{-1}^{1} \int_{-1}^{1} \sin\left(\frac{u}{v}\right) du dv \quad (\text{Much easier to integrate})$$
$$= \frac{1}{2} \int_{-1}^{1} 0 \, dv \quad (\text{Because } \sin\left(\frac{u}{v}\right) \text{ is odd in } u)$$
$$= 0$$

5. Optional Appendix: Why this works



Suppose that D is a small rectangle with sides dx and dy. Then

$$A = \begin{bmatrix} \frac{\partial u}{\partial x} & \frac{\partial u}{\partial y} \\ \frac{\partial v}{\partial x} & \frac{\partial v}{\partial y} \end{bmatrix}$$

transforms D into D', which is an object with sides du and dv:



On the one hand, the area of D' is approximately dudv, but on the other hand, by the formula above:

$$Area(D') = |\det A| Area(D)$$
$$dudv = \left| \det \left[\frac{\frac{\partial u}{\partial x}}{\frac{\partial v}{\partial y}} \frac{\frac{\partial u}{\partial y}}{\frac{\partial v}{\partial x}} \frac{\partial u}{\partial y} \right] \right| dxdy$$
$$dudv = \left| \frac{dudv}{dxdy} \right| dxdy$$

Finally, multiply both sides of the above by f(u, v) = f(x, y) and integrate to get:

$$\int \int_{D'} f(u,v) du dv = \int \int_{D} f(x,y) \left| \frac{du dv}{dx dy} \right| dx dy$$