

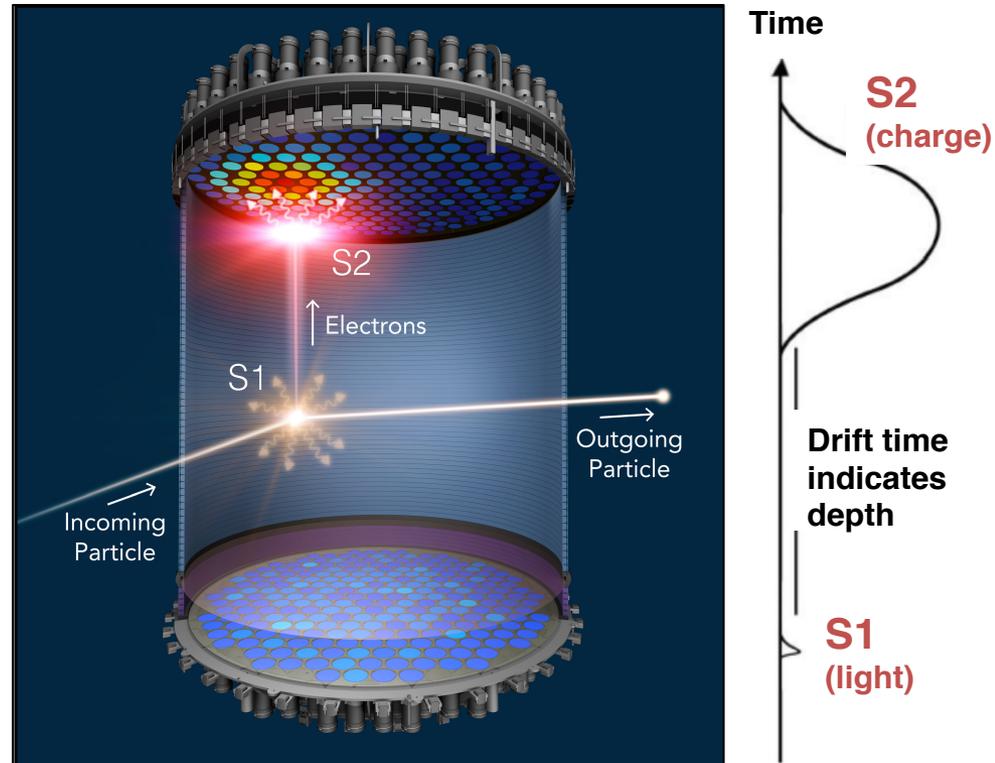
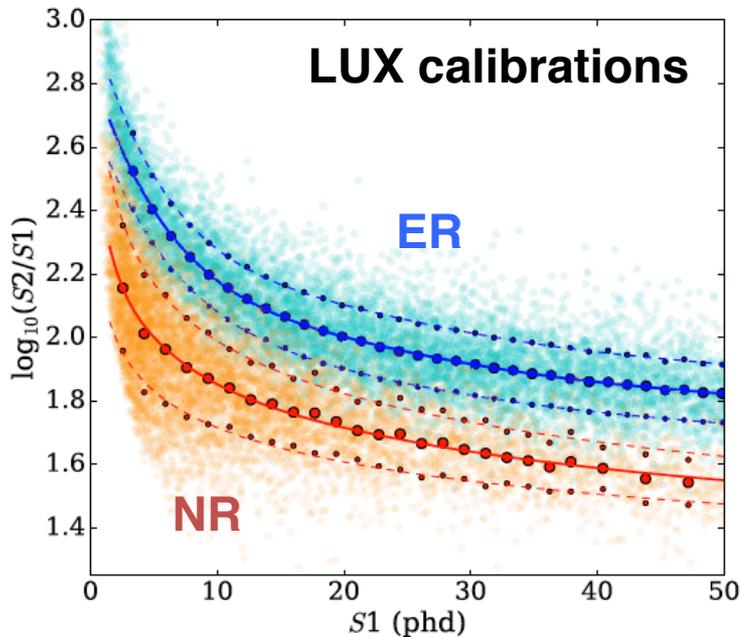


Simulations of external backgrounds at SURF for the LUX and LZ experiments

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On behalf of the LUX and LZ collaborations

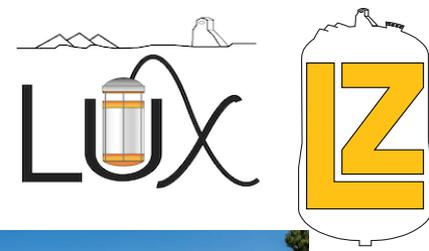
Dual-phase xenon TPC

- ✓ 3D position reconstruction
- ✓ Particle ID: electron recoils (ER) vs nuclear recoils (NR)
- ✓ NR: WIMP signal, neutrons, coherent neutrino scattering
- ✓ ER: Other backgrounds

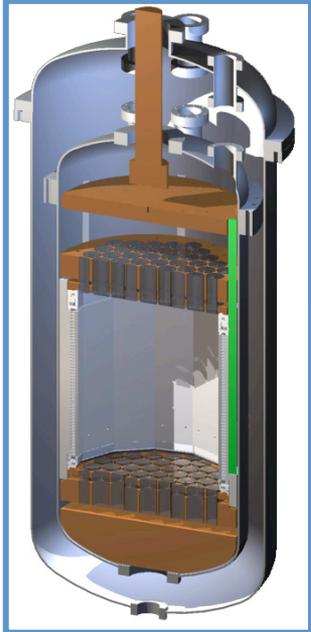


Challenge is to identify very rare ($\ll 1$ event/kg/year) and low energy (< 100 keV) recoils.

LUX and LZ at SURF



LUX

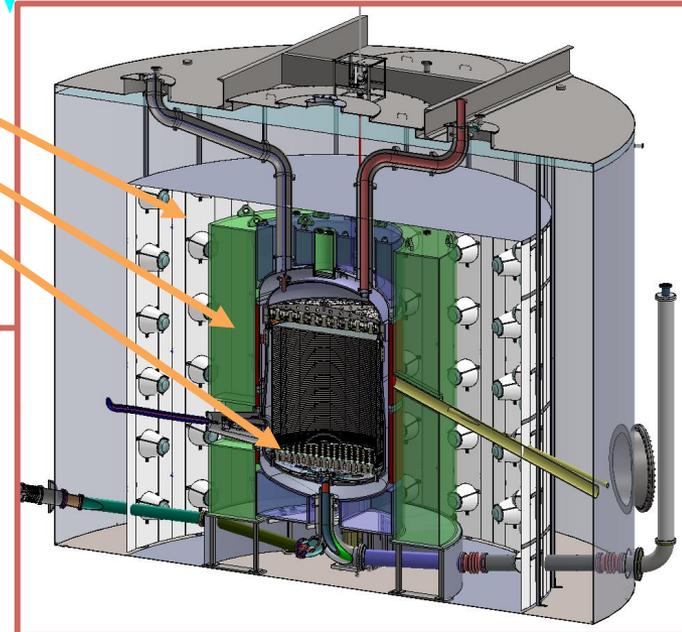


0.5 x 0.5 m dodecagon
370 kg LXe
250 kg active mass
~ 100 kg fiducial mass
122 PMTs

Veto system
Water tank
Scintillator
LXe skin

1.5 x 1.5 m cylinder
10 ton LXe
7 ton active mass
~ 5.6 ton fiducial mass
494 (131) PMTs in TPC (Skin)

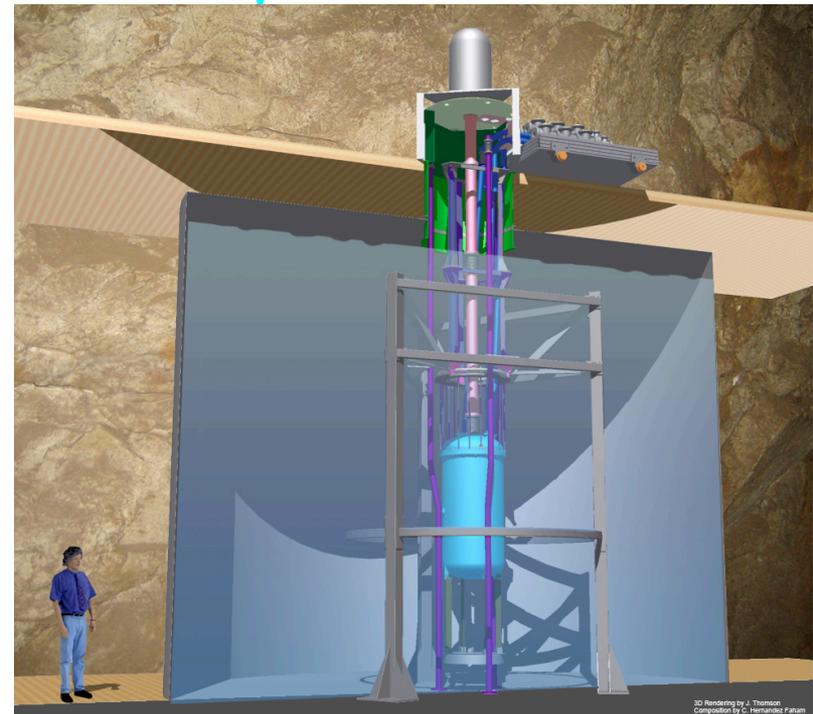
LZ



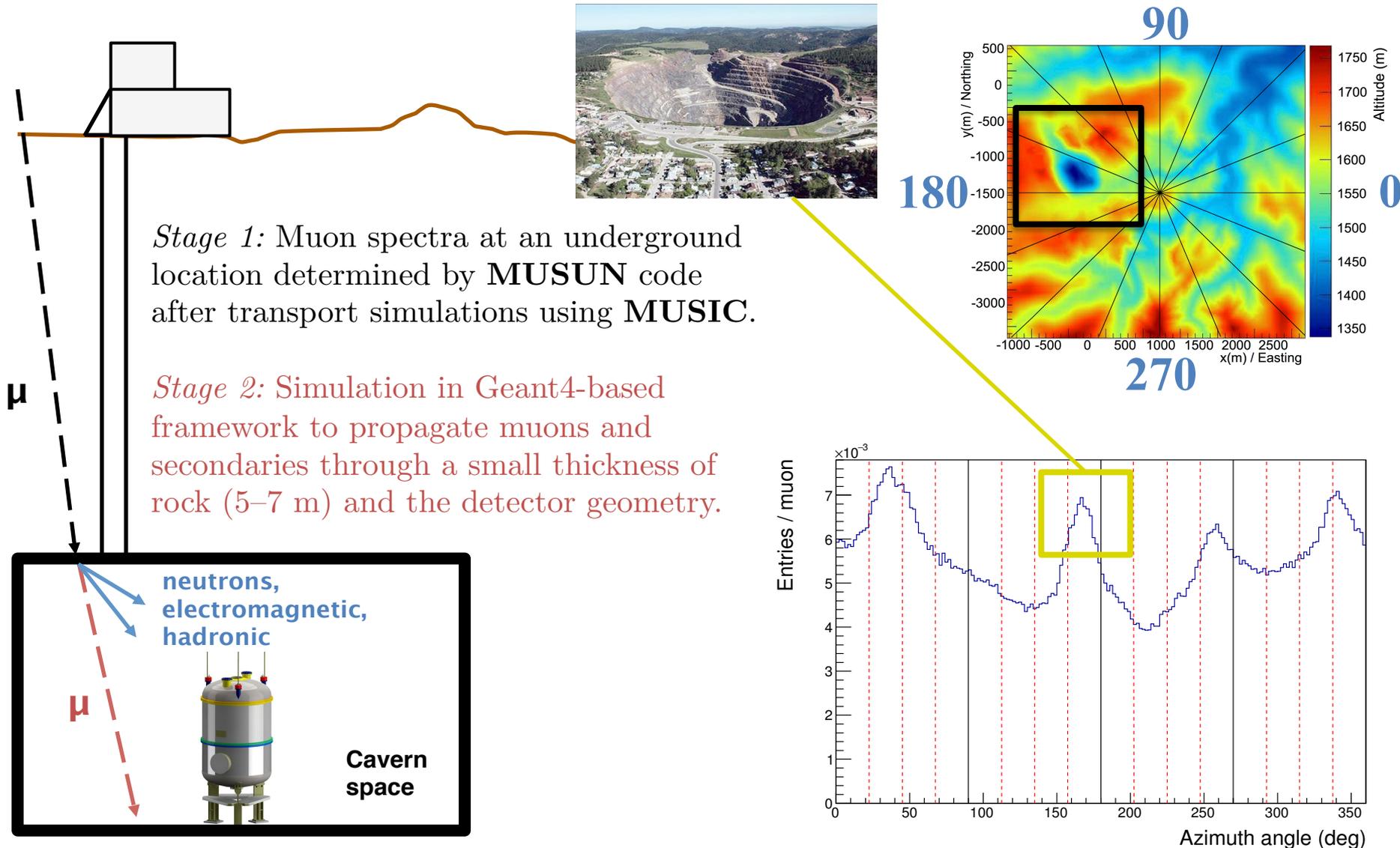
External backgrounds



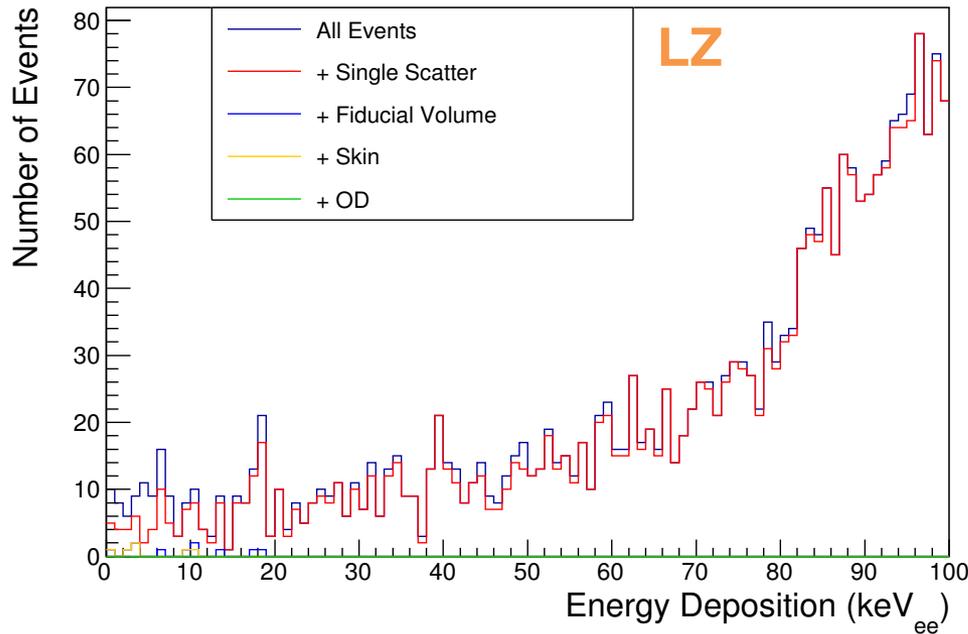
- External background - everything outside of the water tank.
- Here I will talk specifically about muon-induced neutrons and gammas from the cavern rock.



Muon simulations for SURF



Muon-induced background rate



Events that look like WIMPs

- Single scatters: energy-weighted position variance $\sigma_r < 3.0$ cm and $\sigma_z < 0.2$ cm.
- Inside fiducial region: energy-weighted position $r < 68.8$ cm and $1.5 \text{ cm} < z < 132.1$ cm.
- Events without coincident signals (within $500 \mu\text{s}$) above 200 keV_{ee} (100 keV_{ee}) in the scintillator (skin) vetoes.

Cut	Number of surviving events
Region of interest ($< 100 \text{ keV}_{ee}$)	2235
+ single scatter	2128
+ fiducial volume	11
+ skin	6
+ outer detector	0

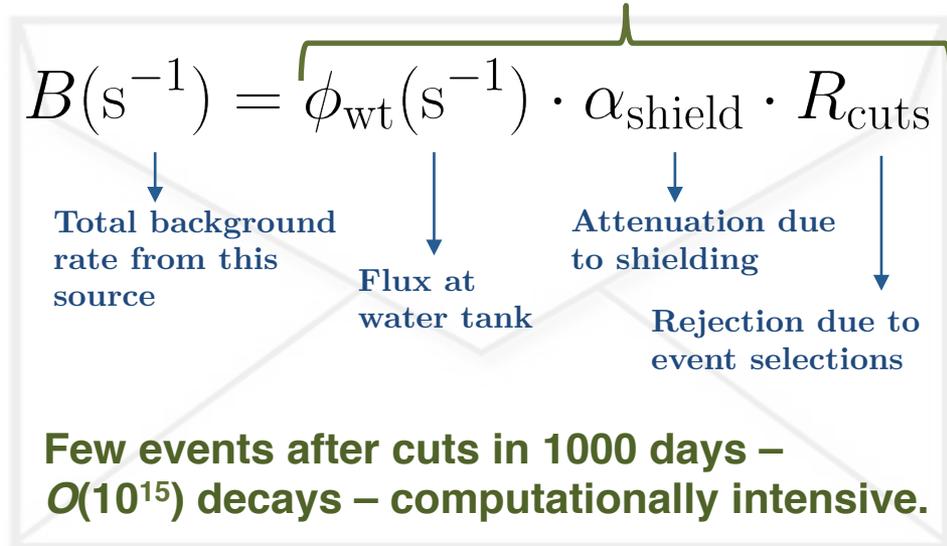
LZ: 2.44 events for the acquired statistics - upper limit of 0.06 events in 1000 days at 90% C.L.

LUX: Re-analysis of LZ simulations with 200 MeV scintillator threshold to mimic water Cherenkov - $(1.0 \pm 0.4) \times 10^{-3}$ events in 300 days.

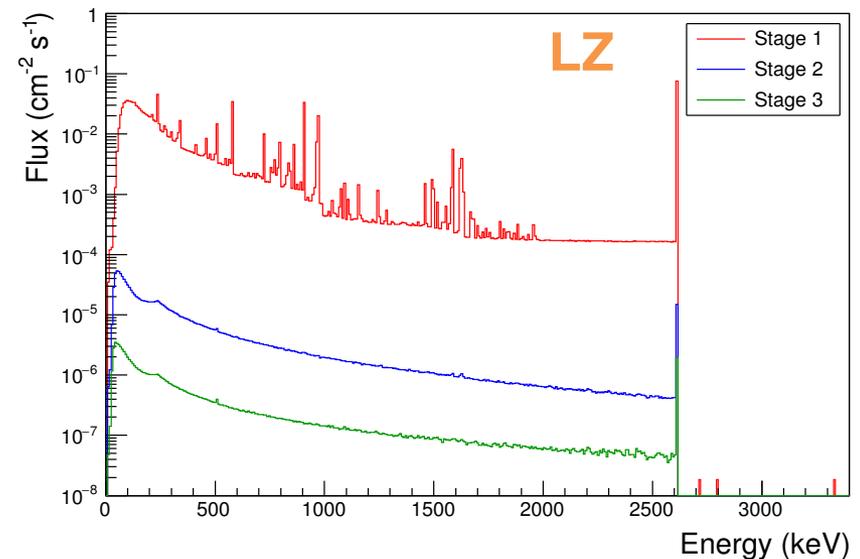
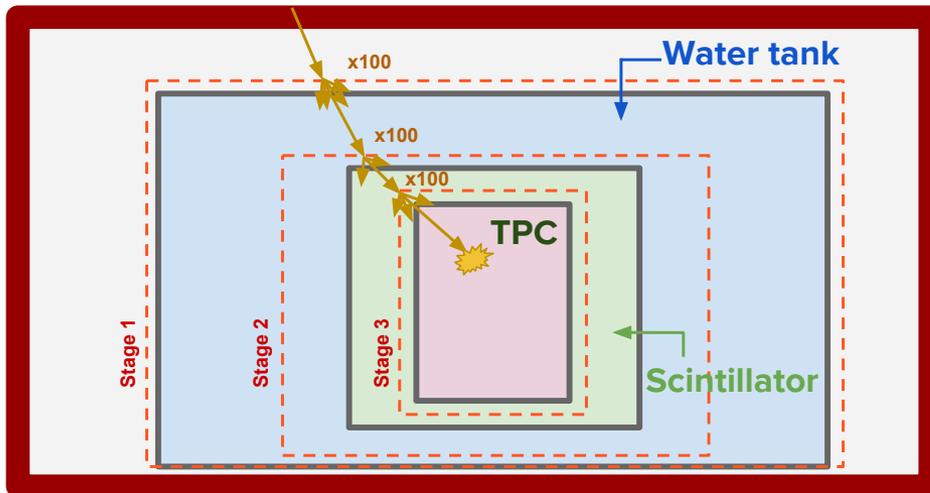
Rock gamma simulations

10^{-9}

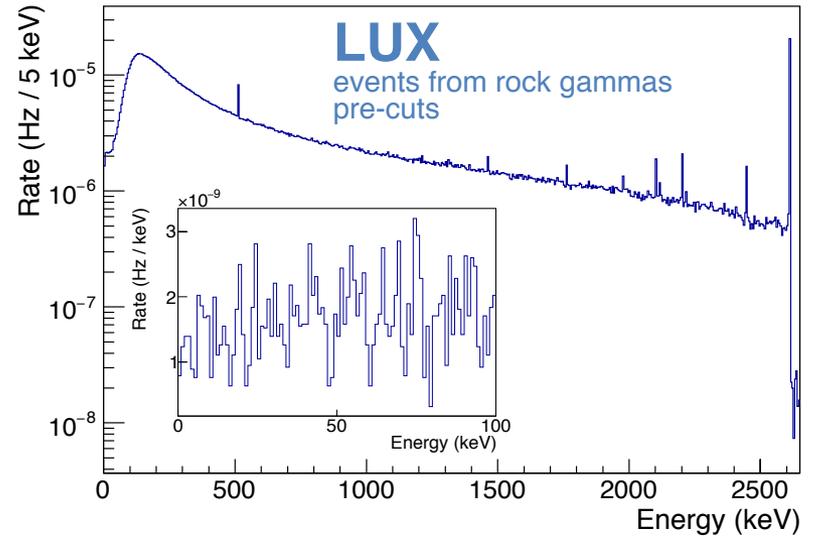
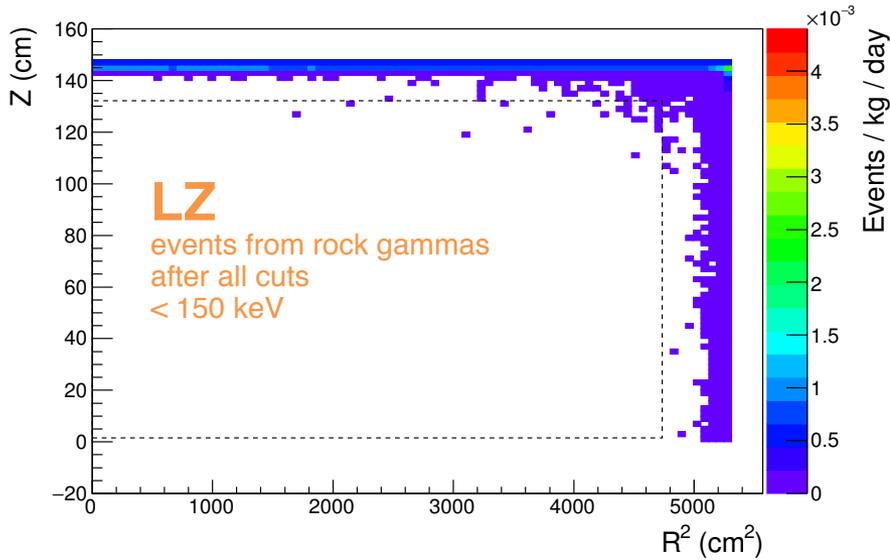
- Start with the decays of ^{238}U , ^{232}Th , ^{40}K in the cavern rock to generate initial γ -ray flux.
- Simulation by brute-force is not possible – improve statistics by using several stages, with a multiplication at each surface.



LZ setup (same principle for LUX simulations)



Rock gamma background rate



Isotope	Activity (Bq/kg)	Simulated live days		Rate (μDRU_{ee})*	
		LZ	LUX	LZ	LUX
^{238}U	24.8	402	7.34×10^4	0.012 ± 0.006	0.23 ± 0.02
^{232}Th	13.6	734	4.02×10^4	0.05 ± 0.02	1.10 ± 0.04
^{40}K	381	262	2.62×10^4	0.014 ± 0.007	0.022 ± 0.009
Total				0.07 ± 0.02	1.35 ± 0.05

* Before S1/S2 discrimination ($> 99\%$)
 $\text{DRU}_{ee} = \text{day}^{-1} \text{kg}^{-1} \text{keV}_{ee}^{-1}$

Conclusions

- External backgrounds are sub-dominant to other sources for both LUX and LZ.

- LZ:**

	ER	NR
Laboratory and Cosmogenics		
Laboratory rock walls	4.6	0.00
Muon induced neutrons	-	0.06
Cosmogenic activation	0.2	-
subtotal	5	0.06
Total	1195	1.03
Total (with 99.5% ER discrimination, 50% NR efficiency)	5.97	0.52
Sum of ER and NR in LZ for 1000 days, 5.6 tonne FV, with all analysis cuts	6.49	

- LUX:**

- Rock gammas: $(1.35 \pm 0.05) \mu\text{DRU}_{ee}$
 - Muon-induced neutrons: $(6.3 \pm 0.3) \mu\text{DRU}_{ee}$
 - Total observed rate: $(3.6 \pm 0.4) \text{mDRU}_{ee}$
- Generators and Monte Carlo techniques can be used for other LUX/LZ studies.